



# Gyro-kinetic simulations of kink instabilities

A. Mishchenko, R. Hatzky, A. Könies, A. Zocco, M. Bordhardt, R. Kleiber

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5.50







5.25 5.00 4.75 4.50 4.25 4.00 3.75 3.50 3.25 3.00 2.75 2.50 2.25 2.00 1.75 1.50 1.25 1.00 0.75 0.50 0.25 0.00 1223:40.950 1223:41.150 122341250 1223:40.800 1223:40.850 1223:40.900 1223:41.000 1223:41.050 1223:41.100 122341200 Coordinated Universal Time - QME-ch02 [keV] - QME-ch04 [keV] - QME-ch06 [keV] \_\_\_QME-ch08 [keV] --- QME-ch11 [keV] --QME-ch13 [keV] - QME-ch15 [keV] ---- QME-ch01 [keV] ---- QME-ch03 [keV] ---- QME-ch05 [keV] ---- QME-ch07 [keV] ---- QME-ch09 [keV] - QME-ch10 [keV] - QME-ch12 [keV] - QME-ch14 [keV]

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- $\bullet$  Current quench events observed in W7-X (OP1.2a/b) related to  $\iota \thickapprox 1$
- A practical issue with limitations of the W7-X operational space
- Likely to be an issue in next stellar ators (not limited to W7-X) ( $\iota \approx 1$  has peculiar optimisation properties [work by H. Weitzner])
- Only parallel current can drive instabilities (diamagnetic current is not destabilising)  $I_{\parallel} \sim 10$  MA in tokamaks vs.  $I_{\parallel} \sim 10$  kA in optimised stellar ators
- Confinement not lost in stellarators even if parallel plasma current disappears
- Transient effect before toroidal current saturates
- Relation of stellarator quenches to sawteeth/NTM problem in tokamaks
- Low shear makes weak effects important: weak ECCD, profiles, turbulence (as in tokamak Neoclassical Tearing Modes, also related to bootstrap)
- Global gyrokinetic description may be needed





- 1995: **GYGLES** is developed at SPC(CRPP) with adiabatic electrons
- $\bullet$  1997: kinetic electrons implemented in GYGLES at IPP
- 1999: ORB5 is developed at SPC(CRPP)
- 1999: EUTERPE is developed at SPC(CRPP)
- 2004: EUTERPE implemented for W7-X at IPP
- 2004: GYGLES becomes electromagnetic at IPP
- 2008: ORB5 becomes electromagnetic, joint development IPP/SPC/UW
- 2009: EUTERPE becomes electromagnetic at IPP

All the codes share the equations solved, physics addressed and the discretisation principles applied. Deeper core routines are often very similar. Normalisation in EUTERPE and ORB5 is almost identical.





- Vlasov equation:  $\partial f/\partial t + \dot{ec{R}} \cdot 
  abla f + \dot{v}_{||} \partial f/\partial v_{||} = 0$
- equations of motion: Shifted Maxwellian to account for  $j_{\parallel 0e} \sim u_{\parallel 0e}$

$$egin{aligned} ec{R} &ec{R} &= \left( v_{\parallel} - rac{q}{m} \langle A_{\parallel} 
angle 
ight) ec{b}^{*} + rac{1}{qB_{\parallel}^{*}} ec{b} imes \left[ \mu 
abla B + q \left( 
abla \langle \phi 
angle - v_{\parallel} 
abla \langle A_{\parallel} 
angle 
ight) 
ight] \ ec{v}_{\parallel} &= \ - rac{1}{m} \left[ \mu 
abla B + q \left( 
abla \langle \phi 
angle - v_{\parallel} 
abla \langle A_{\parallel} 
angle 
ight) 
ight] \cdot ec{b}^{*} \end{aligned}$$

• gyrokinetic quasi-neutrality equation and parallel Ampére's law:

$$egin{aligned} &- 
abla \cdot \left[ \left( \sum_{s= ext{i}, ext{f}} rac{q_s^2 n_s}{k_ ext{B} T_s} 
ho_s^2 
ight) 
abla _ot \phi 
ight] = \sum_{s= ext{i}, ext{e}, ext{f}} q_s \delta n_s \ &\left( \sum_{s= ext{i}, ext{e}, ext{f}} rac{eta_s}{
ho_s^2} - 
abla _ot ^2 
ight) A_ot = \mu_0 \sum_{s= ext{i}, ext{e}, ext{f}} \delta j_{ot s} \end{aligned}$$

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Internal kink mode in straight geometry (screw pinch); MHD regime and FLR regimes; A. Mishchenko and A. Zocco, Phys. Plasmas **19**, 122104 (2012)









- Consider a set of W7-X like iota (double resonance)
- $\bullet$  Consider W7-X like a spect ratio and  $\boldsymbol{\rho}_*$
- Here, flat  $T_i = 2$  keV,  $T_e = 5$  keV
- $\bullet$  Here, flat  $n_0 = 2 \times 10^{19} \text{ m}^{-3}$
- Look for current-driven (1, 1) instabilities
- Study dependence on iota profile





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Magnetic field in screw pinch

$$egin{aligned} ec{B} &= 
abla \psi imes 
abla ec{arphi} + T(\psi) 
abla arphi \ , \ \psi = rac{B_0 r^2}{2} \ , \ T(r) = rac{B_0 r^2}{q r_0} \ ec{j_\parallel} &= rac{1}{\mu_0 r} rac{\mathrm{d}T}{\mathrm{d}r} = rac{B_0}{\mu_0 R_0} \iota(2 - \hat{s}) = rac{B_0}{\mu_0 R_0} (2\iota + r\iota') \ , \ \ \hat{s} = -rac{r\iota'}{\iota} \end{aligned}$$

In W7-X like stellar ators, large part of  $\boldsymbol{\iota}$  is provided externally:

 $\iota = \iota_{\mathrm{ext}} + \iota_{\mathrm{pl}} \ , \ \ \iota_{\mathrm{ext}} \gg \iota_{\mathrm{pl}}$ 

However, the radial derivative of  $\boldsymbol{\iota}$  is mostly provided by internal plasma currents:

$$\iota' = \iota'_{\mathrm{ext}} + \iota'_{\mathrm{pl}} \,, \ \ \iota'_{\mathrm{pl}} \gg \iota'_{\mathrm{ext}}$$

In simulations, we consider for  $s=r_{
m eff}/a$ 

$$u_{
m ext} = 0.88 + 0.12 s^2 \,, \ \ \iota_{
m pl} = 0.12 \exp\left[-rac{(s-0.5)^2}{0.02}
ight] \,, \ \ rac{m{j}_{\parallel} \sim 2 \iota_{
m pl} + r \iota_{
m pl}'}{0.02}$$







$$egin{aligned} &rac{\partial f_1}{\partial t}+\dot{ar{R}}\cdot
abla f_1+\dot{v}_{\parallel}rac{\partial f_1}{\partial v_{\parallel}}=-\dot{v}_{\parallel}^{(1)}rac{\partial f_0}{\partial v_{\parallel}}-\dot{ar{R}}^{(1)}\cdot
abla f_0}{mode\ drive} \ &f_{0e}=n_0\left(rac{2\pi m_e}{T_e}
ight)^{3/2}\exp\left(-rac{m_e v_{\perp}^2}{2T_e}
ight)\exp\left[-rac{m_e (v_{\parallel}-u_0)^2}{2T_e}
ight] \ &u_0=-rac{j_{\parallel 0}}{en_0}=-rac{
ho_{ ext{the}}}{eta_e R_0}(2\iota+r\iota')v_{ ext{the}} \ &
abla f_0=rac{m_e}{T_e}(v_{\parallel}-u_0)
abla u_0f_0&pprox-rac{
ho_{\parallel}}{eta_e R_0}(3\iota'+r\iota'')f_0 \ &v_{ ext{the}}=\sqrt{rac{T_e}{m_e}}\,,\ eta_e=rac{\mu_0n_0T_e}{B_0^2}\,,\ \omega_{ ext{ce}}=rac{eB_0}{m_e}\,,\ 
ho_{ ext{the}}=rac{v_{ ext{the}}}{\omega_{ ext{ce}}}\,,\ 
ho_{\parallel}=rac{v_{\parallel}}{\omega_{ ext{ce}}} \end{aligned}$$





internal  $u_0$  is smaller; internal  $du_0/ds$  and  $df_0/ds$  can be larger (cancellation in  $3\iota' + r\iota''$ )





- Toroidal geometry with W7-X parameters:  $B_0 = 2.5$  T,  $\iota = \iota_{int} + \iota_{ext}$ ,  $j_{\parallel 0} \sim \iota_{int}$  $n_0 = 2. \times 10^{19} \exp \left[-0.67 \tanh \left(\frac{x-0.5}{0.2}\right)\right]$  m<sup>-3</sup>, flat  $T_i = 2$  keV,  $T_e = 5$  keV
- $\bullet$  Global instability is found in the toroidal geometry  $\gamma = 80395.521$  rad/s
- $\bullet$  There is a finite frequency  $\omega = -325859.7405$  rad/s

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effect of  $\nabla n$ : diam. stabilisation?  $\nabla n$  location? effect of  $\nabla T$ ?

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# ORB5 simulations of the "double-kink"



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#### • DONE:

- In [PoP2012]: gyrokinetic internal kink mode in screw-pinch geometry (GYGLES)
- Here: GYGLES simulations of "double kink" for W7-X like iota  $\longrightarrow$  instability found
- Instability becomes stronger with double resonances in iota moving outside
- TO BE DONE:
  - Effect of the temperature and density profiles
  - Effect of bumpiness
  - Effect of toroidicity
  - Effect of fast particles
  - Effect of 3D geometry
  - Gyrokinetic external modes?